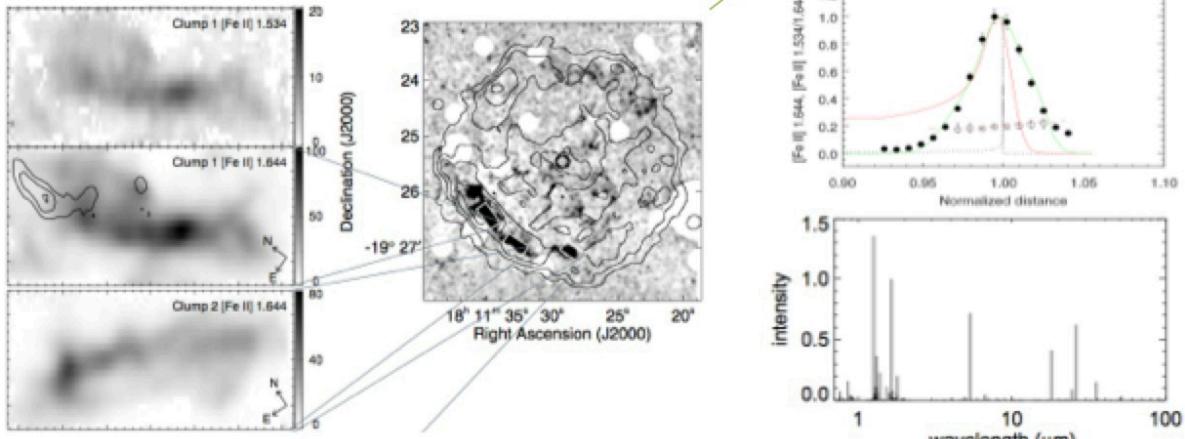


# Radiative SNR shock (Ho-Gyu Lee)

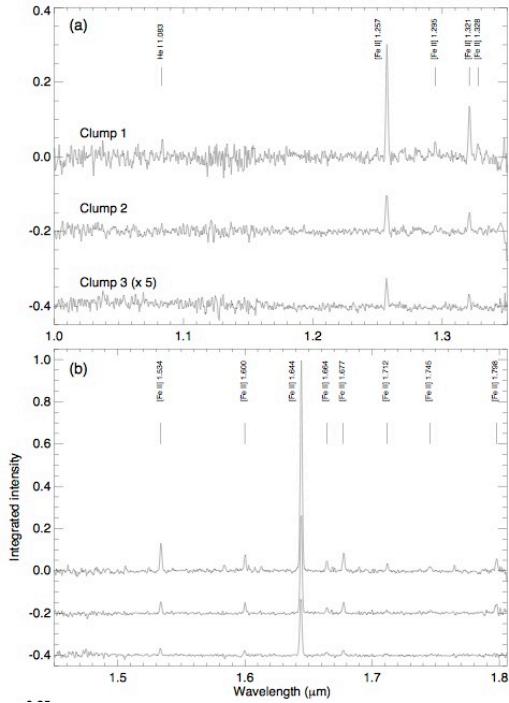
- [Fe II] emission by SNR shock



*Center:* [Fe II] image of G11.2-0.3 with X-ray contours. *Left:* line images constructed by our IFU observations. *Right-top:* comparison of radial profiles with model profiles. Filled circles indicate line intensities and open circles do line ratio. *Right-bottom:* synthesized IR spectrum. (Lee et al. 2013; Koo et al. 2013)

## ❖ G11.2-0.3 (Lee et al. 2013)

- Radial profiles of intensity and line-ratio ( $n_e$ )
  - \* Line intensities obtained at every spaxel using IFU observations
  - \* Dense medium ( $n_o \sim 300 \text{ cm}^{-3}$ ,  $v_s \sim 130 \text{ km/s}$ )
- ⇒ Pre-existing CSM produced by mass-loss.



Integrated spectra of clumps.